

Rainbow metric formalism and Relative Locality

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This proceeding is based on a talk prepared for the XIII Marcell Grossmann meeting. We summarise some results of work in progress in collaboration with Giovanni Amelino-Camelia about momentum dependent (Rainbow) metrics in a Relative Locality framework and we show that this formalism is equivalent to the Hamiltonian formalization of Relative Locality obtained in arXiv:1102.4637.

Keywords: Relative Locality; Rainbow metrics; κ -Poincaré; Curved momentum-space.

I. INTRODUCTION

In literature Planck length¹ $L_P = \sqrt{\hbar G}$ is considered a natural guess as quantum gravity characteristic length scale. However, L_P is non zero only in a regime in which G and \hbar are non zero too, so this hypothesis requires a full fledged quantum gravity theory to elaborate it. We here focus on a "classical non-gravitational" regime [1–3] of quantum gravity. In this regime \hbar and G are both neglected while their ratio is fixed: $M_P = \sqrt{\hbar/G}$. In this paper from now on we will formalize the features of obstruction of measurability and momentum-space deformation with κ -Poincaré Hopf algebra [4, 5] in 1+1D:

$$\{\mathcal{N}, p_0\} = p_i, \quad \{\mathcal{N}, p_1\} = \left(\frac{1 - e^{-2\ell p_0}}{2\ell} - \frac{\ell}{2} p^2 \right), \quad \{p_1, p_0\} = 0, \quad (1)$$

where \mathcal{N} is the boost operator, P_α are the translation operators and where the Casimir operator \mathcal{C} is:

$$\mathcal{C} = \left(\frac{2}{\ell} \sinh \left(\frac{\ell p_0}{2} \right) \right)^2 - p_1^2 e^{\ell p_0}. \quad (2)$$

The algebraic sector of κ -Poincaré can be interpreted as an example of DSR theory[6] with $\ell \sim 1/M_P$ the scale of deformation. It is, therefore, always possible to rely on the existence of a "classical" limit by turning off this deformation ($\ell \rightarrow 0$). In this context moreover it is easier to give a simple interpretation of "Relative Locality" [1, 2] phenomena on physical observables and thus implement a phenomenology [3, 7]. It is well known in literature (see *exempli gratia* [8, 9]) that this kind of phenomenology is rather elusive to theoretical investigation. It would be of paramount importance, then, to formalize some kind of *rainbow metrics* in relative locality, to try to study Planckian effects in curved spacetimes models. In order to do so we need to define the infinitesimal translation of coordinates between two observers labeled by parameters ϵ^α :

$$\delta x^\beta = \epsilon^\alpha \{p_\alpha, x^\beta\}, \quad (3)$$

using Poisson brackets to express the action of translation operators on coordinates, as defined in [1–3].

II. SOME ISSUES WITH RAINBOW METRICS

Smolin and Magueijo [10] proposed that in a framework where free field theories have plane wave solutions, even though the 4-momentum they carry satisfies deformed dispersion relations of the form

$$p_0^2 f^2(p_0/E_P) - (p \cdot p) g^2(p_0/E_P) = m^2, \quad (4)$$

¹ We adopt units such that the speed-of-light scale is 1 ($c = 1$)

spacetime metric should be modified according to the energy of the particle we use to probe it. In fact relation (4) can be realized by the action of a nonlinear map from momentum space to itself, denoted as $U : \mathcal{P} \rightarrow \mathcal{P}$, given by

$$U \cdot (p_0, p_i) = (U_0, U_i) = (f(p_0/E_P) p_0, g(p_0/E_P) p_i),$$

which implies that momentum space has a nonlinear norm, given by $|p|^2 = \eta^{ab} U_a(p) U_b(p)$. If one still wants to have at his disposal a plane wave solution for free fields, since momentum transforms nonlinearly, the contraction between position and momentum must remain linear. Smolin and Magueijo suggested that in case momentum transforms nonlinearly, this can be obtained imposing that $\zeta^{\alpha\gamma} \tilde{g}_{\gamma\beta} = \delta_{\beta}^{\alpha}$, where $\zeta^{\alpha\gamma}$ is the metric of momentum space and $\tilde{g}_{\gamma\beta}$ is the so called "Rainbow metric", such as the spacetime interval

$$ds^2 = \tilde{g}_{\gamma\beta} dx^{\gamma} dx^{\beta} = (dx^0)^2 / f^2 - (dx^i)^2 / g^2, \quad (5)$$

is explicitly energy-dependent. This approach, although extremely valuable from the phenomenologic point of view, is problematic in case one is interested in avoiding to break the line-element invariance. Let's consider, for example the first order expansion of the dispersion relation we can obtain from casimir (2):

$$m^2 = p_0^2 - p_1^2 - \ell p_1^2 p_0. \quad (6)$$

The mass of (6) is clearly invariant under a boost generator of the form

$$\mathcal{N} = x^0 p_1 + x^1 \left(p_0 - \ell p_0^2 - \frac{\ell}{2} p_1^2 \right), \quad (7)$$

obtained imposing $\{\mathcal{N}, \mathcal{C}\} = 0$, assuming $\{p_{\alpha}, x^{\beta}\} = \delta_{\alpha}^{\beta}$. On the other hand, from equation (6) we can set

$$f^2(p_0) = 1, \quad g^2(p_0) = 1 + \ell p_0, \quad (8)$$

and then may identify the element line in the flat spacetime case to be

$$ds^2 = (dx^0)^2 - (1 - \ell p_0)(dx^1)^2, \quad (9)$$

which is not invariant under the ℓ -deformed boost (7).

III. METRIC FORMALISM IN RELATIVE LOCALITY

We will show that is possible to obtain a similar scenario for the κ -Minkowski spacetime framework in which coordinates satisfy the relation $\{\chi^1, \chi^0\} = \ell \chi^1$, and its deformed symmetry generators algebra κ -Poincar time-to-the-right basis defined in Eq. (1). Those relations and operators agree with a deformed symplectic sector:

$$\begin{aligned} \{p_0, \chi^0\} &= 1, & \{p_1, \chi^0\} &= -\ell p_1, \\ \{p_0, \chi^1\} &= 0, & \{p_1, \chi^1\} &= 1. \end{aligned} \quad (10)$$

We can easily obtain the metric formalism as generalization of Pitagora's theorem, expressing the action of translations with Poisson brackets as in Eq.(3):

$$ds^2 \equiv (\epsilon^{\alpha} \pi_{\alpha} \triangleright \xi^{\gamma})(\epsilon^{\beta} \pi_{\beta} \triangleright \xi_{\gamma}) = \{p_{\kappa}, \chi^{\mu}\} \{p_{\lambda}, \chi^{\nu}\} \frac{\partial \pi_{\alpha}}{\partial p_{\kappa}} \frac{\partial \xi^{\gamma}}{\partial \chi^{\mu}} \frac{\partial \pi_{\beta}}{\partial p_{\lambda}} \frac{\partial \xi_{\gamma}}{\partial \chi^{\nu}} d\chi^{\alpha} d\chi^{\beta}, \quad (11)$$

in which (ξ, π) is the set of coordinates used by an observer at rest with respect with the center of mass of a certain process, while (χ, p) is the set used by a generic one. In the $\ell \rightarrow 0$ limit, in which $\{p_{\alpha}, \chi^{\mu}\} \equiv \{p_{\alpha}, x^{\mu}\} = \delta_{\alpha}^{\mu}$, the (11) reduces to

$$ds^2 = \eta_{\alpha\beta} e_{\alpha}^{\gamma} e_{\beta}^{\delta} dx^{\alpha} dx^{\beta} \quad (12)$$

where the tetrads are defined as $e_{\alpha}^{\gamma}(x) \equiv \partial \xi^{\gamma} / \partial x^{\alpha}$. In our case, however we have to take into account the curvature of momentum-space contribution, then $e_{\alpha}^{\gamma}(\chi, p) \equiv \frac{\partial \xi^{\gamma}}{\partial \chi^{\mu}} \frac{\partial \pi_{\alpha}}{\partial p_{\mu}}$. Using (11) it is also possible to find a metric implementing the inferences of momentum-space curvature (also known as Relative Locality effects) on the particle localization process. In fact Eqs. (10) and (11) define a momentum dependent (rainbow) metric, at all orders in ℓ , of the form:

$$\tilde{g}_{\alpha\beta}(\chi, p) = \begin{pmatrix} e_0^0(\chi, p)^2 & -\ell p_1 e_0^0(\chi, p)^2 \\ -\ell p_1 e_0^0(\chi, p)^2 & -(e_1^1(\chi, p)^2 - \ell^2 p_1^2 e_0^0(\chi, p)^2) \end{pmatrix}, \quad (13)$$

which generates an invariant line-element [11]. The Minkowskian limit $\tilde{\eta}_{\alpha\beta}(p)$ of (13) is the result of the inference of a deSitter-like curvature of momentum-space on flat spacetime, formalized with the algebra of symmetries we showed in (1). In this regime, using the geodesic equation for a photon $\tilde{\eta}_{\alpha\beta}\dot{x}^\alpha\dot{x}^\beta = 0$, one can finally find the relation

$$1 - 2\ell p_1 \frac{d\chi^1}{d\chi^0} - (e_1^1(p)^2 - \ell^2(p_1)^2) \left(\frac{d\chi^1}{d\chi^0} \right)^2 = 0 \quad (14)$$

This equation, using the dispersion relation we can obtain from Eq. (2), gives the expression of a worldline (in terms of commutative coordinates x^α) for a massless particle:

$$x^1 - \bar{x}^1 = -e^{\ell p_0}(x^0 - \bar{x}^0), \quad (15)$$

which is the same result defined in Ref. [3, 12] for momentum-dependant massless worldlines. Rainbow metrics formalism is then equivalent to the Hamiltonian one in the spacetime Minkowskian limit of Relative Locality, but it can be even more useful in phenomenology since it naturally implements spacetime curvature. Further analyses should be dedicated to the comparison between this approach and the promising investigation of Relative Locality in curved spacetimes described in [12–14].

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- [1] G. Amelino-Camelia, L. Freidel, J. Kowalski-Glikman, L. Smolin, *The principle of relative locality*, arXiv:1101.0931, Phys. Rev. D **84** (2011), 084010.
 - [2] G. Amelino-Camelia, M. Matassa, F. Mercati, G. Rosati, *Taming nonlocality in theories with Planck-scale-deformed Lorentz symmetry*, arXiv:1006.2126, Phys.Rev.Lett. **106** (2011), 071301.
 - [3] G. Amelino-Camelia, N. Loret, G. Rosati, *Speed of particles and a relativity of locality in κ -Minkowski quantum spacetime* arXiv:1102.4637, Phys. Lett. B **700** (2011) 150-156.
 - [4] S. Majid and H. Ruegg, *Bicrossproduct structure of kappa Poincaré group and noncommutative geometry*, arXiv:hep-th/9405107, Phys. Lett. B **334** (1994) 348.
 - [5] J. Lukierski, H. Ruegg, W. J. Zakrzewski, *Classical quantum mechanics of free kappa relativistic systems*, arXiv:hep-th/9312153, Annals Phys. **243** (1995), 90.
 - [6] G. Amelino-Camelia, *Relativity in space-times with short-distance structure governed by an observer-independent (Planckian) length scale*, arXiv:gr-qc/0012051, Int. J. Mod. Phys. D **11** (2002), 35-60.
 - [7] G. Amelino-Camelia, L. Barcaroli, N. Loret, *Modeling transverse relative locality* arXiv:1107.3334, Int. J. of Theoretical Phys. **51** Issue 11 (2012), 3359-3375.
 - [8] G. Amelino-Camelia, *Are we at the dawn of quantum-gravity phenomenology?*, arXiv:gr-qc/9910089, Lect.Notes Phys. **541** (2000), 1-49.
 - [9] G. Amelino-Camelia, G. Gubitosi, N. Loret, F. Mercati, G. Rosati, *Weakness of accelerator bounds on electron superluminality without a preferred frame*, arXiv:1111.0993, Eur. Phys. Lett. **99** (2012), 21001.
 - [10] J. Magueijo, L. Smolin, *Gravity's Rainbow*, arXiv:gr-qc/0305055, Class. Quant. Grav. **21** (2004), 1725-1736.
 - [11] N. Loret, *Exploring Special Relative Locality with deSitter momentum-space*, arXiv:1404.5093, Phys. Rev. D **90** (2014), 124013.
 - [12] G. Amelino-Camelia, L. Barcaroli, G. Gubitosi, N. Loret, *Dual redshift on Planck-scale-curved momentum spaces*, arXiv:1305.5062, Class. Quantum Grav. **30** (2013), 235002.
 - [13] J. Kowalski-Glikman, G. Rosati, *Relative Locality in Curved Space-time*, arXiv:1303.7216, Mod. Phys. Lett. A **28**, No. 22 (2013), 1350101.
 - [14] G. Amelino-Camelia, L. Barcaroli, G. Gubitosi, S. Liberati, N. Loret, *Realization of DSR-relativistic symmetries in Finsler geometries*, arXiv:1407.8143, Phys.Rev. D **90** (2014), 125030